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## Var 61 Her, Var 62 And, KUV 23012+1702: NEW DWARF NOVAE ON MOSCOW PLATES

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Three new UG-type stars (Var 61, Var 62 and KUV 23012+1702) were discovered in the course of the search for new variables on Moscow archive plates. The coordinates of new dwarf novae, taken from the USNO A1.0 catalogue, are presented in Table 1. The finding charts are shown in Figure 1.

Standard sequences SA62 and SA44 (Priser, 1974) were used to obtain $B$-band magnitudes of comparison stars for $\operatorname{Var} 61$ and $\operatorname{Var} 62$, respectively. Estimates of KUV 23012+1702 were based on the USNO A1.0 catalogue $B$-band scale. The magnitudes of comparison stars are given in Table 2.

All new variable stars are blue on the Palomar prints.
Table 1. Coordinates of New Variables

| Var | $\alpha(\mathrm{J} 2000.0)$ | $\delta(\mathrm{J} 2000.0)$ |
| :--- | :--- | :--- |
| Var 61 | $18^{\mathrm{h}} 05^{\mathrm{m}} 46^{6.4}$ | $+31^{\circ} 40^{\prime} 18^{\prime \prime}$ |
| Var 62 | 001107.3 | +303236 |
| KUV 23012+1702 | 230341.8 | +171755 |

Var 61 Her. The star was estimated on 188 plates taken with the equatorial camera in Moscow and the $40-\mathrm{cm}$ astrograph in Crimea for the interval JD 2415288-49634. Seven outbursts have been observed. The new dwarf nova is bright enough, the range of variability on our plates is $13^{\mathrm{m}} 5-<18^{\mathrm{m}} 0$. The blue magnitude of the star in the USNO A1.0 catalogue is $19^{\mathrm{m}} 3$, so the amplitude of variability, perhaps, is more than $5.8 B$. Outbursts (JD24...):

| \#1 | 30607.197 | $<15.20$ | $\# 3$ | 37842.490 | 14.42 | $\# 6$ | 43803.254 | $<17.00$ |
| :--- | ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: |
|  | 30617.214 | 13.76 |  |  |  |  | 43814.217 | 16.83 |
|  | 30664.122 | $<14.28$ | $\# 4$ | 40014.488 | 15.94 |  | 43815.196 | 17.6 |
| $\# 2$ | 34517.389 | $<15.20$ | $\# 5$ | 43045.244 | 15.20 | $\# 7$ | 45230.283 | $<17.6$ |
|  | 34534.387 | 13.53 |  | 43047.292 | 17.6 |  | 45232.271 | 13.56 |
|  | 34534.405 | 13.58 |  | 43050.290 | 17.8 |  | 45234.249 | 13.46 |
|  |  |  |  |  |  |  | 45251.237 | $<17.6$ |
|  |  |  |  |  |  | 45258.225 | 18.0 |  |



Figure 1. Finding charts


Figure 2. Light curves of KUV 23012+1702: a) and b) bright outbursts; c) faint outburst

Table 2. Comparison Stars

| Var | a | b | c | d | e | f | g | h | i |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Var 61 | 13.15 | 13.63 | 14.28 | 15.20 | 15.65 | 16.80 | 17.00 | 17.6 | 18.0 |
| Var 62 | 15.45 | 15.48 | 15.97 | 16.79 | 17.06 | 17.54 | 17.8 |  |  |
| KUV 23012+1702 | 15.3 | 16.0 | 16.5 | 17.4 | 17.6 | 18.0 |  |  |  |

Var 62 And. The star was investigated on 102 plates taken with the $40-\mathrm{cm}$ astrograph in Crimea (JD 2445266-47835). Five outbursts have been revealed. The range of variability on our plates is $15^{\mathrm{m}} 5-<177^{\mathrm{m}} 8$. Taking into consideration that, in the USNO A1.0 catalogue, the star is shown at $20 .{ }^{\mathrm{m}} 3 B$, we can assume the amplitude of variability exceeding 4 . $8 B$.

Outbursts (JD24...):

| \#1 | 45613.470 | 16.79 | \#2 | 46296.457 | $<17.06$ | \#3 | 47064.427 | $<17.8$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 45614.380 | 16.98 |  | 46303.488 | $<16.79$ |  | 47086.399 | $<15.97$ |
|  | 45615.464 | $<17.8$ |  | 46324.414 | 15.58 |  | 47091.395 | 15.65 |
|  |  |  |  | 46325.480 | 15.55 |  |  |  |
|  |  |  |  | 46327.436 | 15.60 | \# | 47383.475 | 16.63 |
|  |  |  |  | 46329.485 | 15.87 |  | 47389.506 | $<17.54$ |
|  |  |  |  | 46330.412 | 16.46 |  | 47396.502 | $<17.8$ |
|  |  |  |  | 46331.485 | 16.38 |  |  |  |
|  |  |  |  | 46332.457 | > 16.79 | \#5 | 47766.570 | $<16.79$ |
|  |  |  |  | 46352.499 | < 17.06 |  | 47773.551 | 15.45 |
|  |  |  |  | 46357.357 | < 17.8 |  | 47793.472 | $<17.54$ |

KUV 23012+1702. The star was firstly discovered by Kondo et al. (1984), as a new blue variable object. No classification is given in their work.

I independently discovered variability and estimated this variable on 156 plates taken with the $40-\mathrm{cm}$ astrograph in Crimea (JD2444076-47477). Sixteen outbursts have been revealed. The cycle is (very approximately) 27 days. The range of variability on our plates is $15^{\mathrm{m}} 8-<18 \mathrm{~m}^{\mathrm{m}} 0$. It would be interesting to know the brightness in minimum. In the USNO A1.0 catalogue, the object is shown at $18.2 B$, but it is uncertain (because of short cycle) if this really shows the star at minimum light.

The new variable is a good candidate to UGSU-type dwarf novae. Two kinds of outbursts were found: bright ones have $15 \mathrm{~m} .8 B$ in maximum and a duration of more than 8 days (Figures 2a and 2 b ); faint ones have $16^{\mathrm{m}} 5 B$ in maximum and a duration less than 5 days (Figure 2c).

A CCD spectrum of KUV 23012+1702 was obtained by Wegner and Dupuis (1993). According to their work, the spectral type of this object is sdBe, that does not contradict the suggested classification.

Outbursts (JD24...):

| \#1 | 44487.415 | 16.68 | \#7 | 45614.307 | < 18.0 | \#12 | 46058.219 | 15.76 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 44488.497 | 16.68 |  | 45615.355 | 16.48 |  |  |  |
|  | 44492.409 | < 17.6 |  | 45615.389 | 16.59 | \#13 | 46289.464 | $<17.4$ |
|  |  |  |  | 45616.339 | 17.00 |  | 46293.430 | 16.68 |
| \#2 | 44852.427 | < 17.6 |  | 45616.373 | 17.04 |  | 46295.430 | $<17.4$ |
|  | 44854.461 | 17.04 |  | 45617.314 | 18.0 |  |  |  |
|  |  |  |  | 45617.376 | 17.6 | \#14 | 46677.439 | 16.50 |
| \#3 | 45228.475 | 16.20 |  | 45619.310 | < 18.0 |  | 46679.424 | 16.91 |
|  | 45232.444 | < 18.0 |  |  |  |  | 46681.425 | < 18.0 |
|  |  |  | \#8 | 45642.333 | 17.13 |  |  |  |
| \#4 | 45530.443 | 15.82 |  | 45644.278 | 17.68 | \#15 | 47035.492 | 16.82 |
|  | 45531.463 | 16.10 |  | 45646.301 | $<18.0$ |  | 47041.485 | < 17.6 |
|  | 45536.497 | 16.77 |  |  |  |  |  |  |
|  |  |  | \#9 | 45695.204 | 16.00 | \#16 | 47382.431 | 15.76 |
| \#5 | 45553.518 | $<17.4$ |  |  |  |  | 47386.363 | 15.93 |
|  | 45558.381 | 17.27 | \#10 | 45936.458 | 15.79 |  | 47390.407 | 16.64 |
|  | 45560.494 | < 17.4 |  | 45943.455 | 16.82 |  | 47396.420 | < 18.0 |
|  | 45562.487 | < 18.0 |  | 45947.458 | $<17.6$ |  |  |  |
| \#6 | 45580.403 | 17.13 | \#11 | 45964.417 | $<18.0$ |  |  |  |
|  | 45581.537 | 18.0 |  | 45965.348 | < 17.6 |  |  |  |
|  | 45582.499 | < 18.0 |  | 45972.427 | 16.73 |  |  |  |

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